SPIN PRECESSION IN A 2-BODY SYSTEM: A NEW TEST OF GENERAL RELATIVITY

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- 1. Newtonian Theory
- 2. General Relativistic Corrections
 - a. Periastron Precession
 - b. Spin Precession
- 3. Relativistic Spin Precession in the Double Pulsar [Breton et al., *Science* 321, 104 (2008)]
- 4. Double Pulsar Numbers

(Spin Precession ~5°/yr. etc.)

- 5. Chart Comparing Various 1 & 2 Body Systems
- 6. GPB and Other Tests
- 7. Theoretical Approach Outline

NEWTONIAN (1687/KEPLER (1619)



FIGURE 5-12. Kepler's law of equal areas: the two shaded areas are equal, so the speed of the planet is greater near perihelion because it must sweep out a larger arc of the ellipse in the time $t_2 - t_1$ compared to the equal time $t_2' - t_1'$.

$F = \frac{Gm_1m_2}{r^2} \Rightarrow \text{Closed Orbit (Eclipse)}$

 \vec{L} = Orbital Angular Momentum $\rightarrow \bot$ to orbit \vec{A} = Runge - Lenz Vector

→ From focus to perihelion

$$\rightarrow \perp$$
 to \vec{L}





GENERAL RELATIVITY: 1-BODY/NO SPIN





$$\xrightarrow{MERCURY-SUN} \approx 3 \times \left(3 \times 10^{-8}\right) \left(8.3 \times 10^{-7}\right) rads / s = 43"/CENTURY$$

2. GENERAL RELATIVISTIC CORRECTIONS

DEF: a = Semi - Major Axis; e = Eccentricity $M = m_1 + m_2$ $\mu = \frac{m_1 m_2}{M}$ $\vec{L} = L\vec{n}(\vec{n} = \hat{L}) = \text{Angular Momentum}$ $\vec{A} = \text{Runge} - \text{Lenz Vector}$

$$\mathcal{O}_{g} = \frac{GM}{c^{2}a} = \text{Gravitational Coupling Constant} = \frac{(\overline{\omega}a)^{2}}{c^{2}} = \left(\frac{v_{av}^{2}}{c^{2}}\right)$$

$$\boxed{O} = \frac{2\pi}{T} = \left(\frac{GM}{a^3}\right)^{1/2} = \frac{L/\mu}{a^2(1-e^2)^{1/2}} = \text{Average Orbital Angular Velocity}$$
$$(T = \text{Period})$$

ALL precessions per unit time $\sim \alpha_g \overline{\omega}$

(a) Periastron precession (Robertson/EIH)

 $\frac{d\vec{A}}{dt} = \left(\vec{\Omega}^{*(E)} \times \vec{A}\right) = \left(\frac{d\omega}{dt}\right) (\vec{n} \times \vec{A}) [\omega = \text{Argument of Perihelion}]$

$$\frac{d\omega}{dt} = \frac{3}{1 - e^2} \left(\alpha_g \overline{\omega} \right)$$

(b) Spin Precession (Barker-O'Connell)

$$\frac{d\vec{S}^{(1)}}{dt} = \left(\vec{\Omega}_{so}^{(1)} \times \vec{S}^{(1)}\right) + \text{Spin} - \text{Spin term (smaller)}$$
$$= \left|\vec{\Omega}_{so}^{(1)}\right| \left(\vec{n} \times \vec{S}^{(1)}\right)$$

GENERAL CHECKS

1.
$$\frac{d\vec{A}}{dt} = \vec{\Omega}^* \times \vec{A}$$
$$\frac{d\vec{L}}{dt} = \vec{\Omega}^* \times \vec{L}$$

Hence

$$\frac{d}{dt} \left(\vec{A} \cdot \vec{L} \right) = 0$$

 $\Rightarrow \vec{L}$ is always \perp to \vec{A} , as is required

GENERAL CHECKS

2.
$$\frac{d\vec{J}}{dt} = \frac{d}{dt} \left\{ \vec{L} + \vec{S}^{(1)} + \vec{S}^{(2)} \right\} = 0$$

Conservation of total momentum

Science 321, 104 (July 4, 1008)

Relativistic Spin Precession in the Double Pulsar Breton et al.,

measure the relativistic precession of pulsar B's spin axis

4.77°/yr observed 5.07°/yr theory [4°.78 for A] \downarrow \downarrow \downarrow PERIODS \rightarrow 71 yrs <u>75 yrs</u>

B. M. Barker and R. F. O'Connell, *Phys. Rev. D* **12**, 329 (1975)

"predicted by general relativity"



Figure 1: Schematic view of the double pulsar system showing the important parameters for the modeling of pulsar A's eclipse (dimensions and angles are not to scale). Pulsar B is located at the origin of the cartsian coordinate system, whereas the projected orbital motion of <u>pulsar A during its eclipse</u> is parallel to the y axis at a constant z_0 as seen from Earth, which is located toward the positive *x* axis. Because the orbital inclination is almost perfectly edge-on (14), we can approximate the z axis to be coincident with the orbital angular momentum. The spin axis of pulsar B, whose spatial orientation is described by θ and ϕ , is represented by the Ω vector. The magnetic axis of pulsar B corresponds to the μ vector and makes an angle with respect to Ω . Lastly, the absorbing region of the diplar magnetosphere of pulsar B, truncated at radius R_{mag} , is shown as a shaded red region.

10. DOUBLE PULSAR SYSTEM: THE WORKS

Science 303, 1143 (Feb. 20, 2004)



A is \approx 3,600 times as energetic as B!! Two-Pulsar dance. Schematic of the double pulsar system (not to scale) relative to observers on Earth. The ellipses are the orbits of the two pulsars A and B around the common center of gravity seen at an oblique angle.

Pulsar A's strong outflow of relativistic particles and magnetic fields ("pulsar wind")penetrates into the light cylinder of star B and causes formation of a bow shock with long tail behind pulsar B. The light cylinder (with radius l_c) plays an essential role in the generation of the radio beams that cause the observed pulsed signal.

The beam of pulsar B is depicted here as a hollow cone centered on the magnetic pipole axis. The distruption of pulsar B's light cylinder on the side facing pulsar A may short-circuit the currents in B's magnetosphere that produce the radio beams, which might explain the weakness of the pulses of B observed over most of its orbital cycle.

Changes in orientation of the light cylinder will cause <u>variations in the emitted beam</u>, as will <u>relativistic precession of the rotation axis</u>.

 $l_c = \frac{c}{\Omega^{spin}} \rightarrow$ co-rotation of charged particles in magnetosphere cannot persist beyond the surface where the tangential velocity = c

4. DOUBLE BINARY NUMBERS

<u>Distance to Earth</u> 2.2×10¹⁶ miles ≈ $\left(\frac{1}{22}\right)$ Milky Way Diameter Number of Pulsars ≈1,700 $R_{\Theta} \approx 7 \times 10^{2} km$ n - Star Radius = 10 km $\approx 10^{-5} R_{\Theta}$

 \rightarrow only 3 in nearby galaxies

Binary with 1 Pulsar and 1 Neutron star 6

Double Pulsar 1 (Very Unique!)

 $a = \text{Semi-Major Axis} = 1.25 R_{\Theta} = 8.7 \times 10^5 km$

 \Rightarrow Entire Binary could fit within our Sun

e = Eccentricity = 0.088

 $i = Angle of Inclination = 88.69^{\circ}$

 \Rightarrow System is observed nearly perfectly edge - on

T = Orbital Period = 2.45 Hours

Pulsar	Mass	Pulse Period	Pulse Frequency	Surface B field (G)	Magneto- sphere Extent	***** Spin Precession	Time for 360° Revolution of Spin
$A(m_2)$	1.3381	22.7	$44 s^{-1}$	6×10^{9}	1,084 <i>km</i>	4°.782	75 years
$B(m_1)$	1.2489	2773.5	$0.36s^{-1}$	2×10^{12}	132,000 <i>km</i>	5°.0734	71 years

Periastron Precession $\left(\frac{d\omega}{dt}\right) = 16.89949(68) / \text{Year} \approx (43,000) \text{ PP of } Hg$

 $\approx m_1 + m_2 = 2.58708(16) \rightarrow \text{High Accuracy}$

Spin Precession $| \rightarrow$ need separate values for m_1 and m_2

$$\rightarrow \frac{m_1}{m_2}$$
 is obtainable from $a_1 = \frac{m_2}{M}a$ and $a_2 = \frac{m_1}{M}a$

and the measured projected semi-major values $\left(\frac{a_1 \sin i}{c}\right)$ and $\left(\frac{a_2 \sin i}{c}\right)$

 $\rightarrow m_1$ and m_2 from **Shapiro Time Delay** [delay of 6.2 μs due to A]

 \rightarrow 13% Accuracy with improvements very likely

EMISSION

 $B \approx 10^8 G - 10^{10} G$

 \Rightarrow Radio Emission

BIRTH

Supernova Explosions of Stars $\approx (16-18)M_{\Theta} \quad [(4-10)M_{\Theta} \quad \text{for solitary pulsars}]$

TABLE COMPARING VARIOUS SYSTEMS

$$\alpha_g = \frac{GM}{c^2 a} \text{ where } M = m_1 + m_2; \ a = \text{Semi Major Axis}$$
$$\overline{\omega} = \frac{2\pi}{T} = \left(\frac{GM}{a^3}\right)^{1/2} \text{ where } T = \text{ Orbital Period}$$
$$1^\circ/\text{yr} = \left(3 \cdot 6 \times 10^2\right) \text{ arc - sec/yr}$$
$$1 \text{ rad/sec} = \left(1 \cdot 8 \times 10^9\right)^0/\text{yr} = \left(6 \cdot 5 \times 10^{12}\right) \text{ arc - sec/yr}$$

System	Sun-Mercury	Earth-Gyro (GP-B)	PSR 1913+16	* PSR * J0737-3039A/B (Double Binary)	Sun- (Earth/Moon) Gyro
$\alpha_{_g}$	3×10^{-8}	7×10^{-10}	2×10^{-6}	4.4×10^{-6}	9.8×10^{-9}
$\overline{\omega}(s^{-1})$	8.3×10^{-7}	10 ⁻³	2.2×10^{-4}	$7 \times 10^{-4} s^{-1}$	2×10^{-7}
$\alpha_{g}\overline{\omega}(s^{-1})$	2.5×10^{-14}	7.5×10^{-13}	4.4×10^{-10}	$3.1 \times 10^{-9} s^{-1}$	1.9×10^{-15}
$\alpha_{g}\overline{\omega}(^{\circ}/\mathrm{yr})$	$\left(4.5\times10^{-5}\right)^{\circ}/\mathrm{yr}$	$(1.4 \times 10^{-3})^{\circ}/\mathrm{yr}$	(0.79)°/yr	(5.6)°/yr	12 <i>mas</i> /yr
Periastron Precession (P.P.) $\frac{3\alpha_g\overline{\omega}}{1-e^2}$	43"/century	(12.6)"/yr	4°.2/yr	16°.9/yr	36 <i>mas</i> /yr
Spin-Orbit Precession $\approx \frac{3}{2} (\alpha_g \overline{\omega}) \approx \frac{1}{2}$ (P.P.)	21"/century [irrelevant due to strong tidal interaction with sun]	ntury nt due $6.6"/yr$ $1^{\circ}.1/y$ g tidal[Spin-Spin] is $0.04"/yr$]		5°.1/yr (Pulsar B) 4°.8/yr (Pulsar A	19 <i>mas</i> /yr
а	$6 \times 10^7 kms$	$6 \times 10^3 kms$	$2 \times 10^6 kms$	$9 \times 10^5 kms$ = 1.25 R_{Θ}	$1.5 \times 10^8 km$
M/M_{Θ}	1	3×10^{-6}	1.44 (Pulsar) 1.39 (Companion)	$1.34(A) \rightarrow m_2$ $1.25(B) \rightarrow m_1$ $M_A / M_B = 1.07$	1
e	0.206	0.0014	0.6	0.08	0.017

OBSERVATIONS $\left[1mas = \left(2 \cdot 8 \times 10^{-7}\right)^{\circ}\right]$

i) <u>Single Pulsar</u> \rightarrow Stairs, et al., *PRL* <u>93</u>, 141101 ('04)/0.51°/yr (95% confidence)/2-Body

- ii) <u>Lense-Thirring effect on 2 lageos satellites [spin (earth)-earth orbit]</u>
 Sattellite-Laser-Ranging (precession ≈ few mms.) → Ciufolini and Paulis,
 Nature 431, 958 ('04)/ 31 (and 31.5) mas/yr / Nodal (orbital) Motion of satellite
- iii) Lunar Laser Ranging [Spin (earth-moon)-Sun orbit]
 - → JPL etc., PRL 98, 071102 ('07)
 - \rightarrow 0.1% Accuracy/Nodal motion of moon

iv) <u>Double Pulsar</u> \rightarrow Breton et al., Science 321, 104 (2008) \rightarrow 5.07%/yr (13% precision) [Agrees with B-O'C, *PRD* <u>12</u>, 329 ('75)]

v) <u>Gravity Probe B</u> → gyro orbiting earth
 → NASA Final Report, Dec. 2008
 Spin-Orbit (Geodesic) → 6,614 · 4mas/year (error 0 · 5%)
 Spin-Spin (Frame Dragging) → 39mas/year (error 15%) [170 times smaller]