



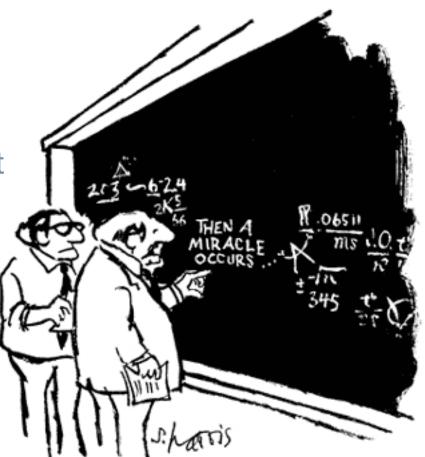


Sagnac effect in general relativity and Gödel's Universe

Endre Kajari, Wolfgang P. Schleich, First LARES workshop, Sapienza, Universita di Roma, July 4th, 2009

Outline of the talk:

- Gödel's Universe
- Sagnac's original experiment and its description in GR
- Sagnac effect in a proper reference frame (PRF)
- Double eight-loop
 interferometer (DELI)



"I think you should be more explicit here in step two."

S. Harris

Why did Gödel bother himself with rotating universes?



Page 3

- In Sept. 1946 Einstein received a letter by Gamov
- Einsteins response: "What does it mean, that the Universe as a whole possesses an angular momentum?"
- It is very likely that Einstein discussed the suggestion of Gamov with Gödel
- May 7th, 1949: Gödel's lecture on rotating universes at the Institute for Advanced Studies

Gödel's Universe

Line element in Gödel's Universe

$$ds^{2} = c^{2}dt^{2} - \frac{dr^{2}}{1 + \left(\frac{r}{2a}\right)^{2}} - r^{2}\left(1 - \left(\frac{r}{2a}\right)^{2}\right)d\phi^{2} - dz^{2} + 2r^{2}\frac{c}{\sqrt{2}a}dtd\phi$$

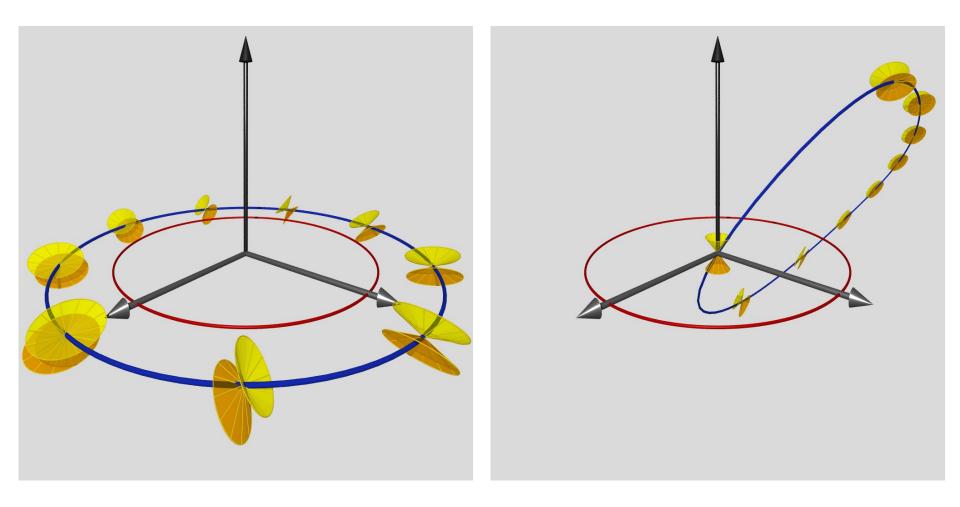
In the limit of large critical Gödel radius: flat spacetime

$$ds^2 \stackrel{a \to \infty}{\longrightarrow} c^2 dt^2 - dr^2 - r^2 d\phi^2 - dz^2$$

Gödel's Universe is (i) spacetime homogeneous (ii) stationary (iii) anisotropic (iv) rotational symmetric with a constant rotation vector along z-axis

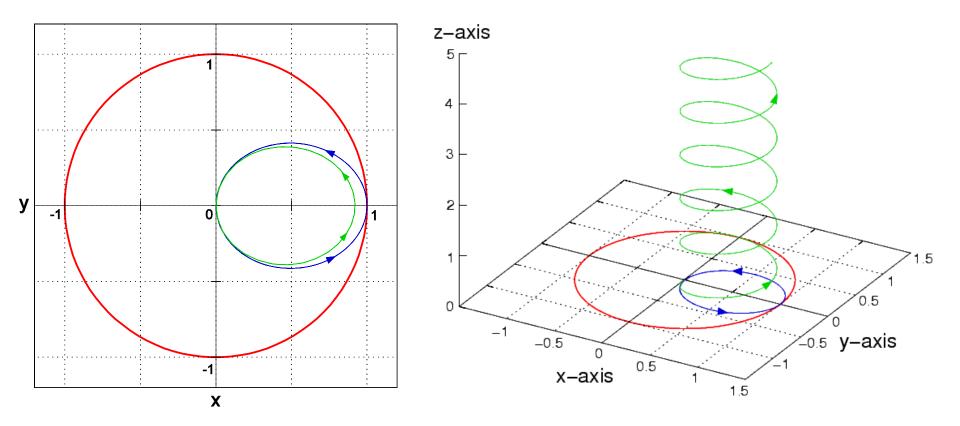
Closed Timelike Worldlines in Gödel's Universe y х

Closed Timelike Worldlines in Gödel's Universe

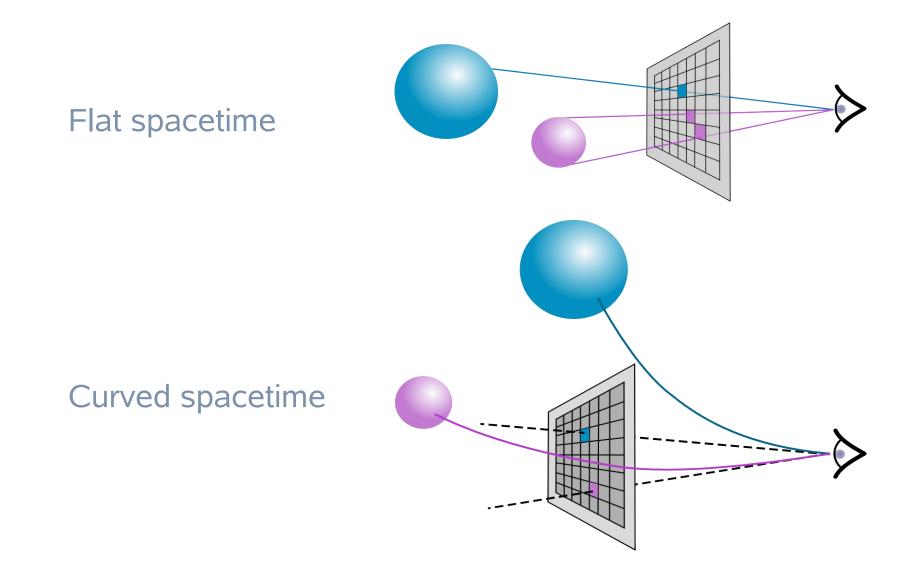


Page 6

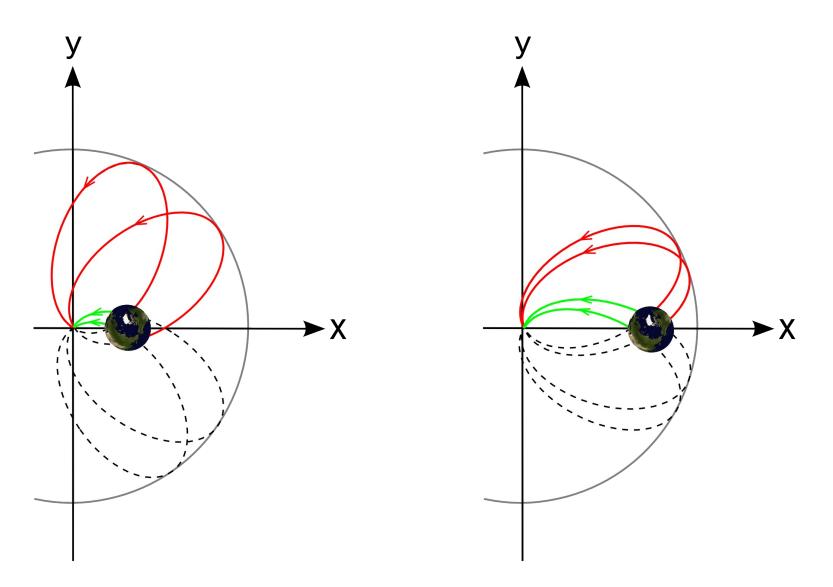
Null-geodesics starting at the origin



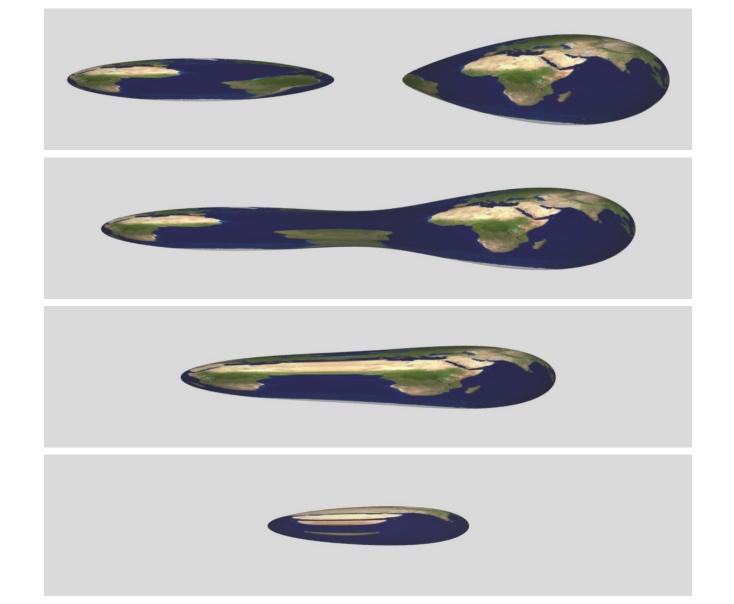
Visualization in General Relativity



View on Earth in Gödel's Universe



View on Earth in Gödel's Universe





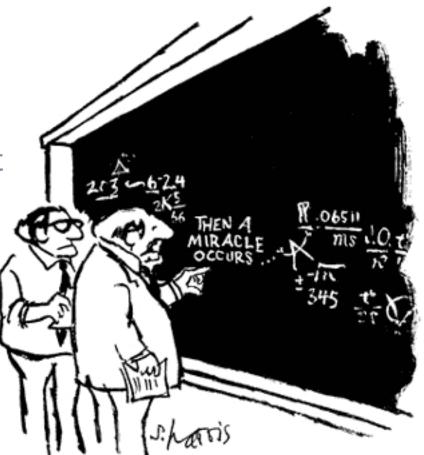
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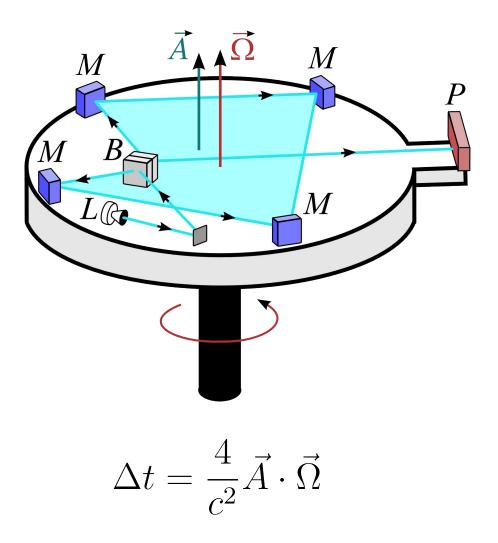
Double eight-loop interferometer (DELI)



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Sagnac's Original Experiment



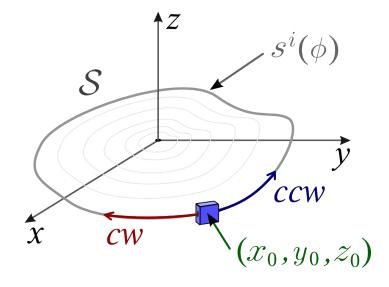
Sagnac's Conclusion:

"The observed interference effect is clearly the optical whirling effect due to the movement of the system in relation to the ether and directly manifests the existence of the ether, supporting necessarily the light waves of Huygens and of Fresnel."

C. R. Acad. Sci. 157, 708 and 1410 (1913), translated by R. Hazelett

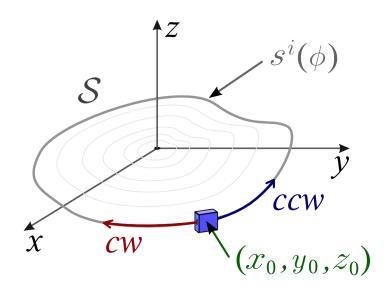
Sagnac Time Delay in General Relativity

for a time independent metric

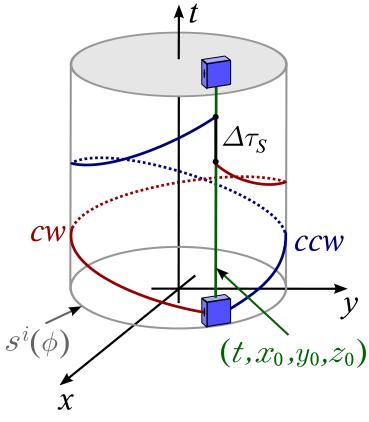


Sagnac Time Delay in General Relativity

for a time independent metric

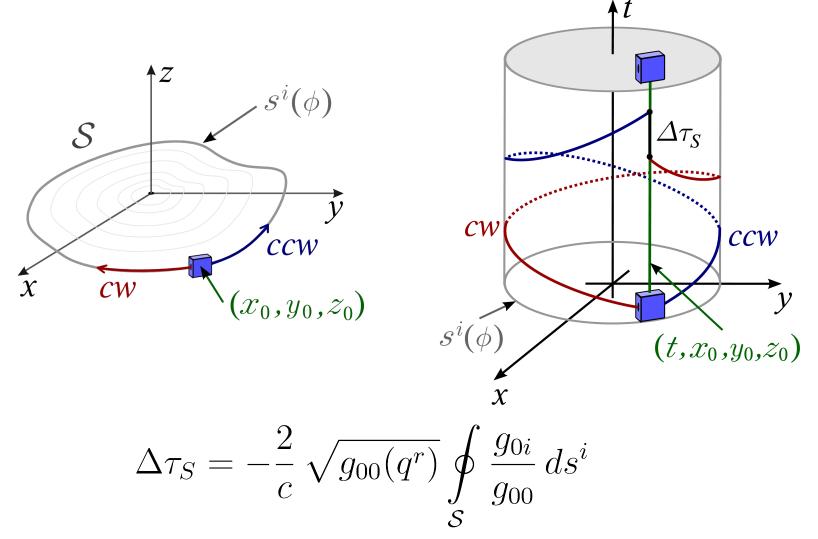


Page 15

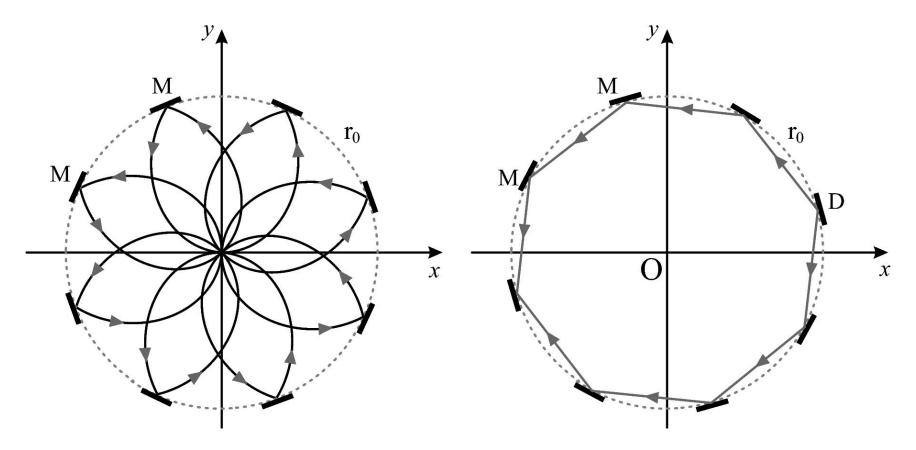


Sagnac Time Delay in General Relativity

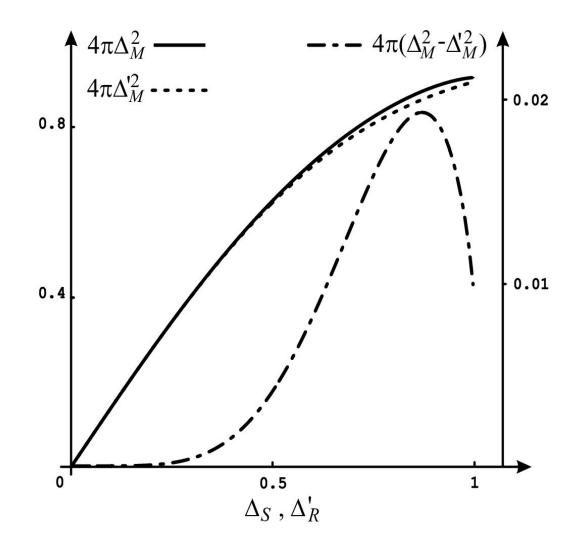
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Motivation



Motivation



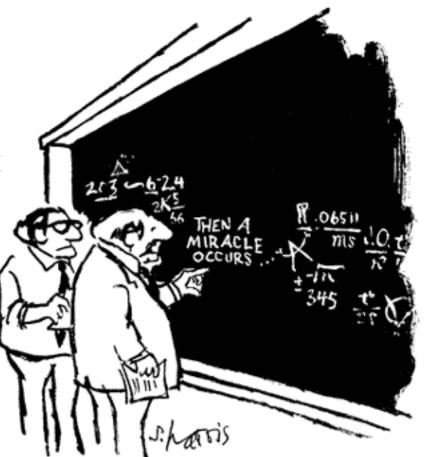
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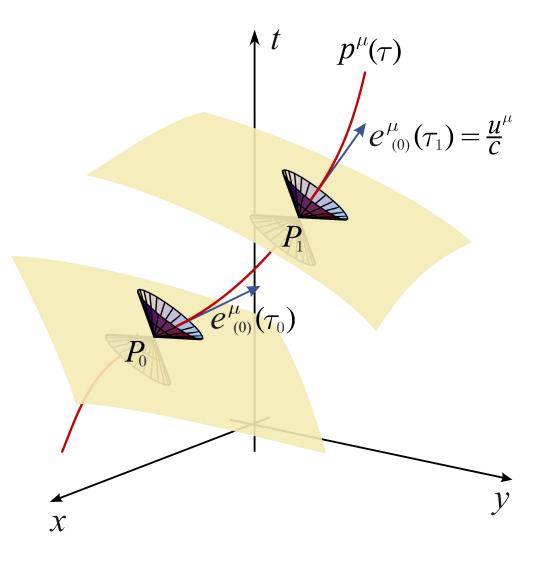
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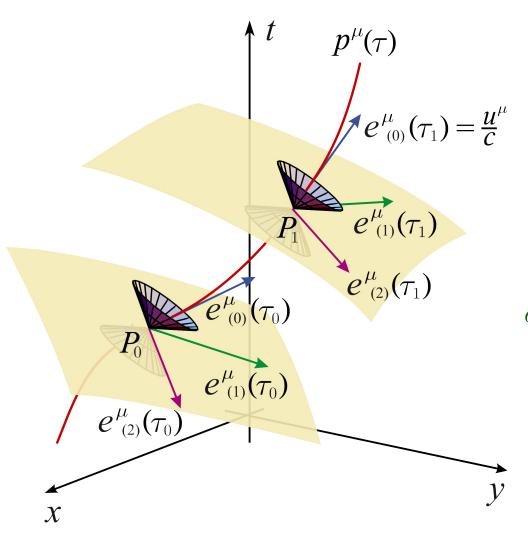
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four-velocity: $u^{\mu}(\tau) = \frac{dp^{\mu}}{d\tau}$

four-acceleration:

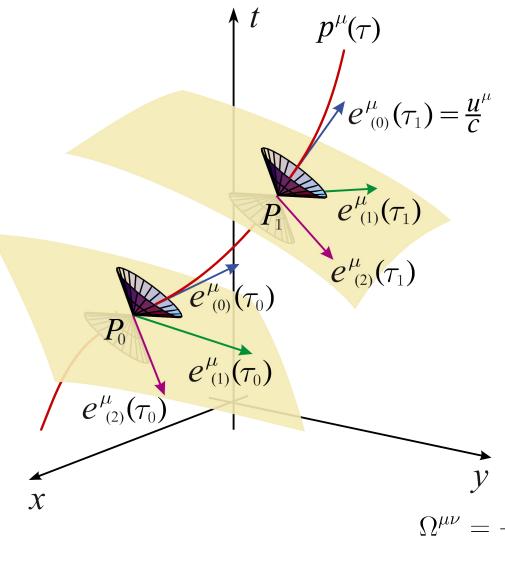
 $a^{\mu}(\tau) = u^{\mu}_{;\nu} u^{\nu}$



four-velocity: $u^{\mu}(\tau) = \frac{dp^{\mu}}{d\tau}$

four-acceleration: $a^{\mu}(\tau) = u^{\mu}_{;\nu} u^{\nu}$

orthonormal tetrads: $e^{\mu}_{(\alpha)}(\tau) e^{\nu}_{(\beta)}(\tau) g_{\mu\nu}(p^{\sigma}(\tau)) = \eta_{(\alpha\beta)}$



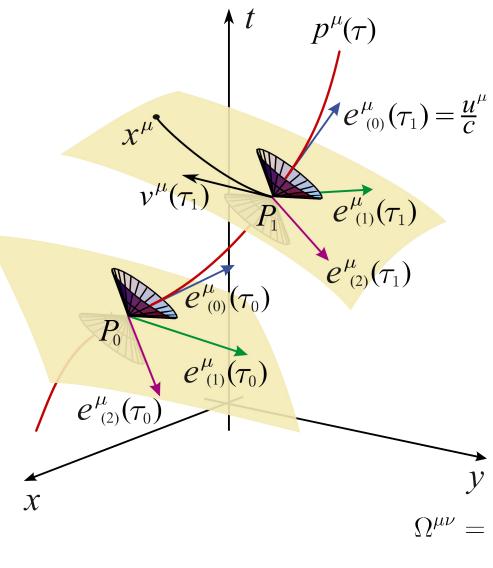
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orthonormal tetrads: $e^{\mu}_{(\alpha)}(\tau) e^{\nu}_{(\beta)}(\tau) g_{\mu\nu}(p^{\sigma}(\tau)) = \eta_{(\alpha\beta)}$

proper transport: $e^{\mu}_{(\alpha);\nu} u^{\nu} = -\Omega^{\mu}_{\ \nu} e^{\nu}_{(\alpha)}$ transport matrix:

 $\Omega^{\mu\nu} = -\frac{1}{c^2} \left(a^{\mu} u^{\nu} - a^{\nu} u^{\mu} \right) + \frac{1}{c} u_{\rho} \omega_{\sigma} \varepsilon^{\rho\sigma\mu\nu}$



four-velocity: $u^{\mu}(\tau) = \frac{dp^{\mu}}{d\tau}$

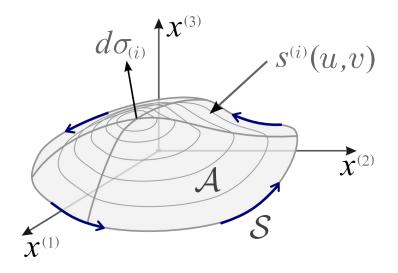
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Sagnac Time Delay in a PRF

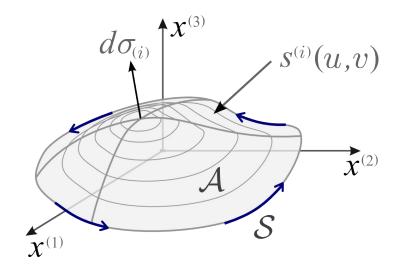


general surface

$$A_{(a)} = \iint\limits_{\mathcal{A}} d\sigma_{(a)}$$

$$A_{(a)}^{(i_1)} = \iint\limits_{\mathcal{A}} s^{(i_1)} d\sigma_{(a)}$$

Sagnac Time Delay in a PRF



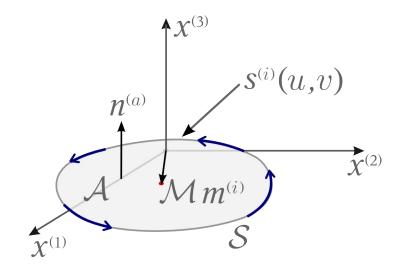
general surface

$$A_{(a)} = \iint\limits_{\mathcal{A}} d\sigma_{(a)}$$

$$A_{(a)}^{(i_1)} = \iint_{\mathcal{A}} s^{(i_1)} d\sigma_{(a)}$$

$$\Delta \tau_{S} = \frac{4}{c^{2}} \sqrt{g_{(00)}(q^{(r)})} \left[-\omega^{(a)} A_{(a)} + \frac{2c}{3} \varepsilon^{(0ajk)} R_{(0)\{(i_{1})(j)\}(k)}(p^{(r)}) A_{(a)}^{(i_{1})} + \frac{1}{c^{2}} \left(\omega^{(l)} a_{(l)} \delta_{(i_{1})}^{(a)} - 3 \omega^{(a)} a_{(i_{1})} \right) A_{(a)}^{(i_{1})} + \mathcal{O}\left(A_{(a)}^{(i_{1})(i_{2})}\right) \right]$$

Sagnac Time Delay in a PRF



planar surface

$$A_{(a)} = \mathcal{A} n_{(a)}$$

$$A_{(a)}^{(i_1)} = \mathcal{M} \, m^{(i_1)} \, n_{(a)}$$

$$\begin{aligned} \Delta \tau_S &= \frac{4}{c^2} \sqrt{g_{(00)}(q^{(r)})} \left[-\omega^{(a)} A_{(a)} + \frac{2c}{3} \varepsilon^{(0ajk)} R_{(0)\{(i_1)(j)\}(k)}(p^{(r)}) A_{(a)}^{(i_1)} \right. \\ &+ \frac{1}{c^2} \left(\omega^{(l)} a_{(l)} \delta_{(i_1)}^{(a)} - 3 \, \omega^{(a)} a_{(i_1)} \right) A_{(a)}^{(i_1)} + \mathcal{O}\left(A_{(a)}^{(i_1)(i_2)} \right) \right] \end{aligned}$$

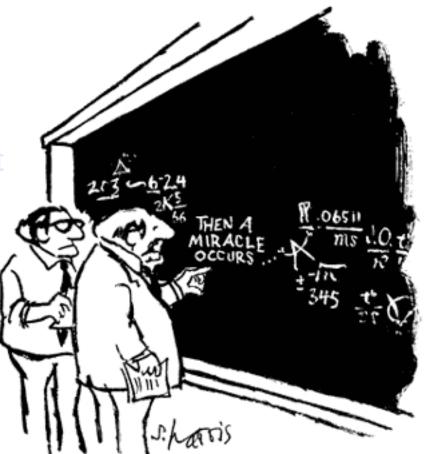
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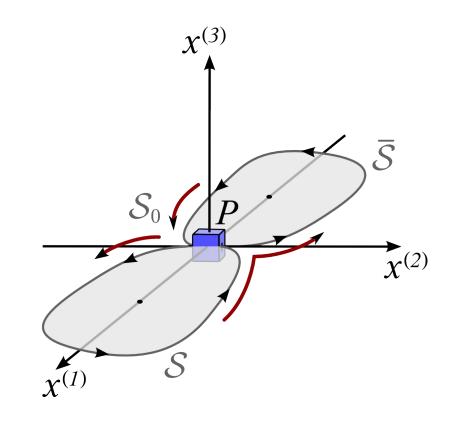
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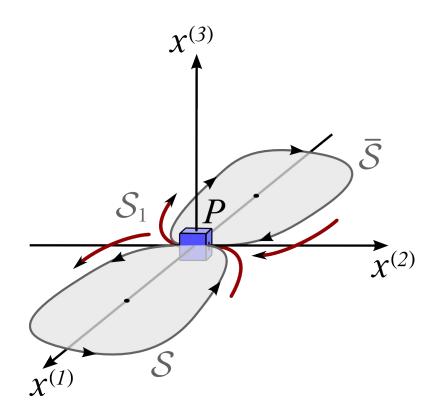
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DELI: Measurement mode 1



$$\Delta \tau_{Sp}(\mathcal{S}_0) = -\frac{4}{c^2} \,\omega^{(a)} n_{(a)} \cdot 2\mathcal{A} + \mathcal{O}\left(A_{(a)}^{(i_1)(i_2)}\right)$$

DELI: Measurement mode 2



$$\Delta \tau_{Sp}(\mathcal{S}_1) = n_{(a)} \left[\frac{8}{3c} \varepsilon^{(0ajk)} R_{(0)\{(i_1)(j)\}(k)}(p^{(r)}) + \frac{4}{c^4} \left(\omega^{(l)} a_{(l)} \delta^{(a)}_{(i_1)} - 3\omega^{(a)} a_{(i_1)} \right) \right] m^{(i_1)} \cdot 2\mathcal{M}$$

Sagnac effect in a rotating reference frame (flat spacetime) and in Gödel's spacetime

	DELI operation mode 1	DELI operation mode 2
Flat spacetime Δau_{Sp}	$\frac{4}{c^2} \mathcal{N}\Omega_R n^{(3)} \cdot 2\mathcal{A}$	$\frac{12}{c^4} \mathcal{N}^3 r_0' \Omega_R^3 n^{(3)} m^{(1)} \cdot 2\mathcal{M}$
Gödel's spacetim Δau_{Sp}	$-rac{4}{c^2}\Omega_G\;n^{(3)}\cdot 2\mathcal{A}$	0
with constants:	$\Omega_G = \frac{c}{\sqrt{2}a}$	$\mathcal{N} = rac{1}{\sqrt{1 - \left(rac{r_0'\Omega_R}{c} ight)^2}}$

The team (Institute of Quantum Physics, Ulm University)



Michael Buser



Cornelia Feiler



Wolfgang P. Schleich

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Cornelia Feiler



Wolfgang P. Schleich

E. Kajari, M. Buser, C. Feiler and W. P. Schleich, "Rotation in relativity and the propagation of light" in the Proceedings of the International School of Physics "Enrico Fermi" Course "Atom Optics and Space Physics" (2009)

Thank you very much

for your attention!

