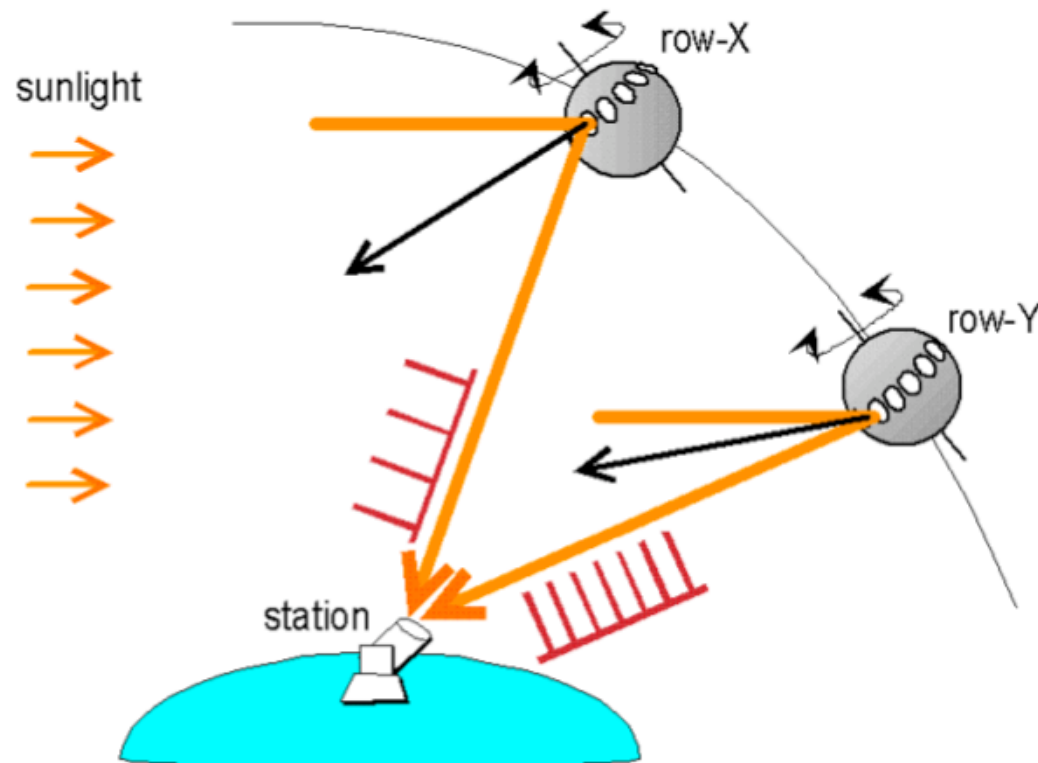


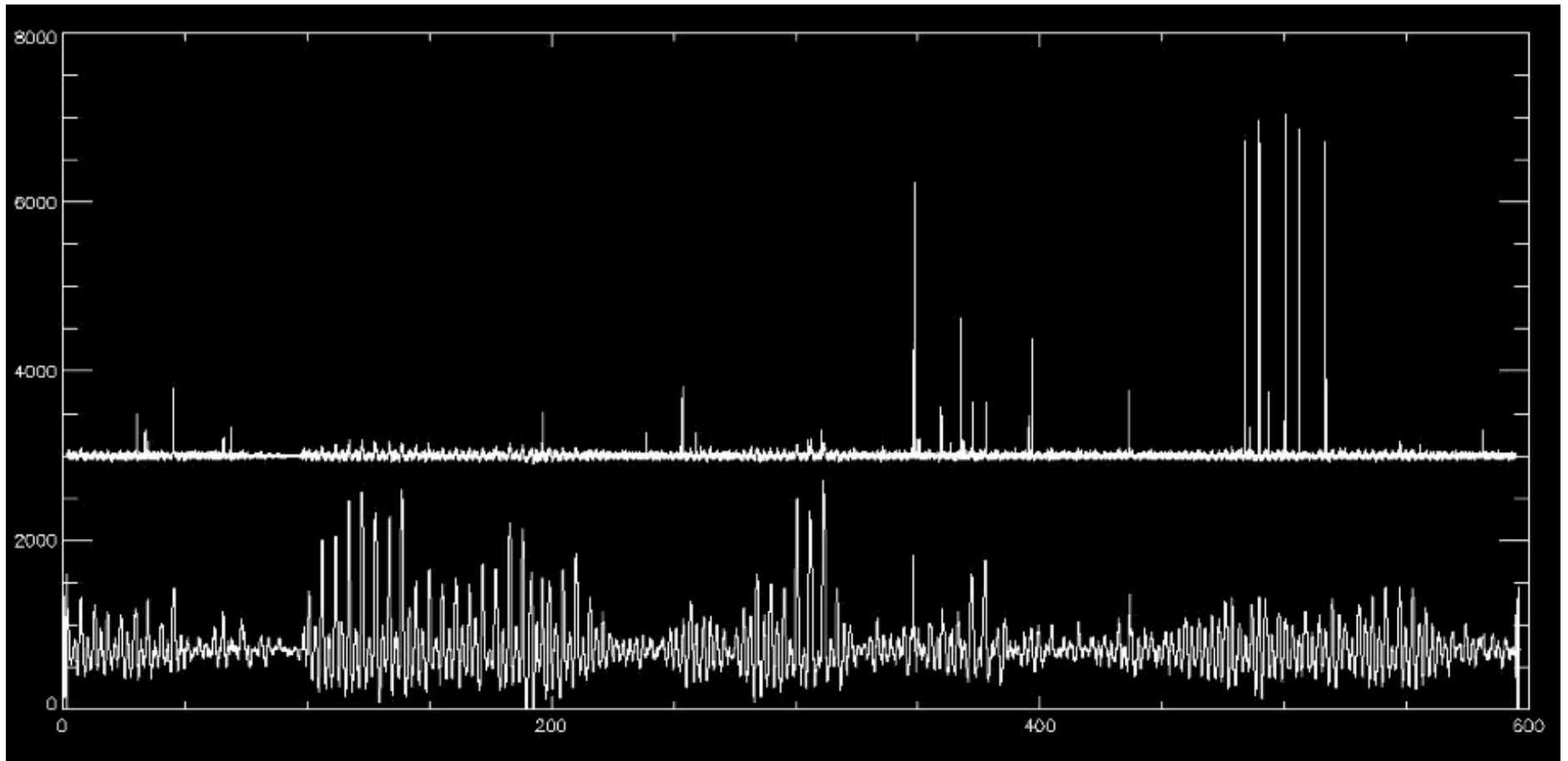
# Glints



**Figure 2.** *How solar glints occur. Note that, at different points in the orbit, Sunlight is reflected towards the telescope from different rows; and that rows with fewer reflectors give rise to longer intervals between glints.*

R. Wood , T. Otsubo and R. Sherwood *LAGEOS 2 Spin Rate and Orientation*  
[http://cddisa.gsfc.nasa.gov/lw13/docs/papers/target\\_wood\\_1m.pdf](http://cddisa.gsfc.nasa.gov/lw13/docs/papers/target_wood_1m.pdf).

# RULLI Observations of LAGEOS II



## 3.6 meter telescope observations

Currie, Dell'Agnello, Dell Monache, Tauraso (2007)

# Ajisai Spin Parameter Determination Using Graz Kilohertz Satellite Laser Ranging Data

- Georg Kirchner, Walter Hausleitner, and Elena Cristea

*IEEE Transactions on Geoscience and Remote Sensing* **45**, no. 1, (January 2007), 201 (2kHz laser, Graz Ausria)